

About the existence of a Poincaré transition in rotating nuclei

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Breaking of left-right symmetry -

a reminder of a famous publication

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Equilibrium Configurations of Rotating Charged or Gravitating Liquid Masses with Surface Tension. II*

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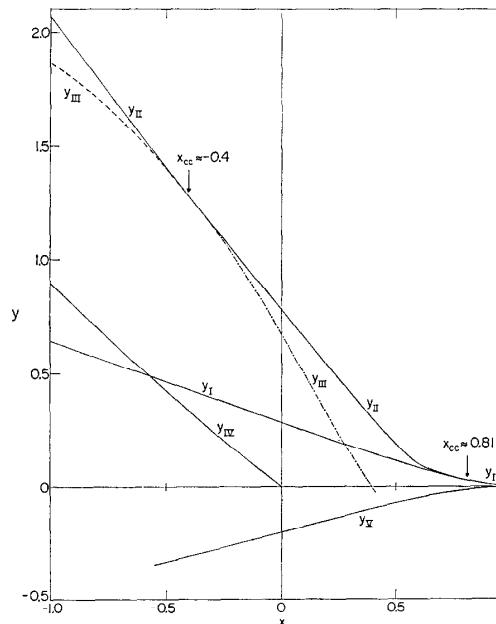


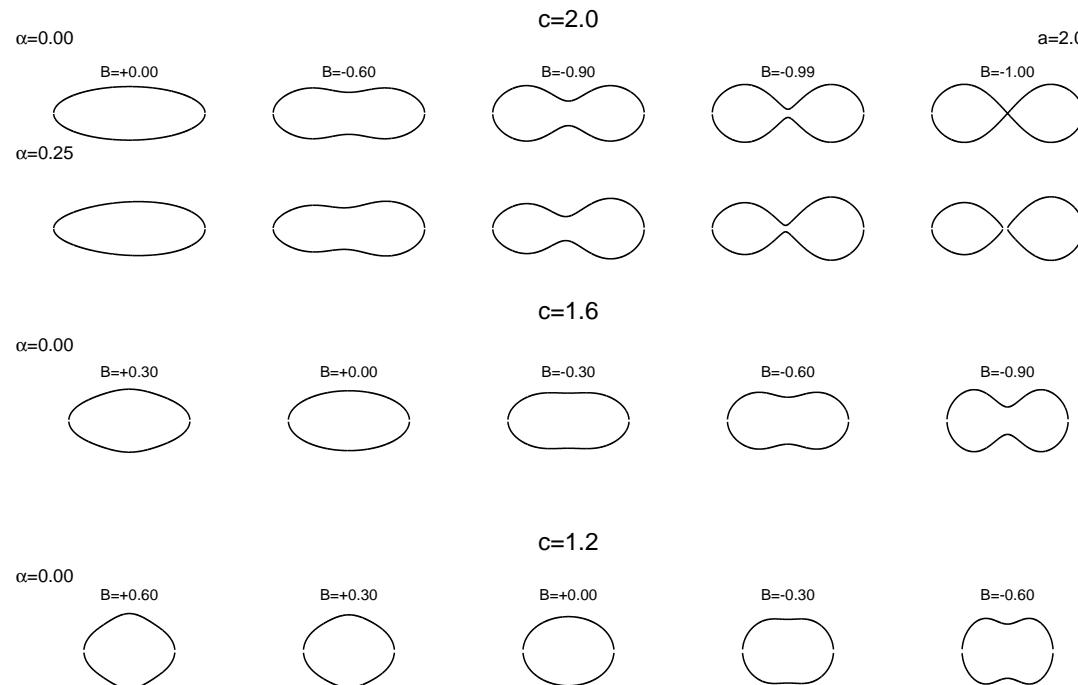
FIG. 2a. Various critical rotational parameters y in their dependence on the fissility parameter x . Triaxial shapes appear between y_I and y_{II} . Saddle shapes are stable against reflection asymmetric distortions to the right of the dot-dashed portion of y_{III} . Triaxial shapes are unstable against asymmetry between the dashed portion of y_{III} and y_{II} . The critical curves y_{IV} and y_V will be discussed in future installments of this series of papers.

Modified Funny-Hills Shape parametrisation

- Huge variety of nuclear shapes (g.s. \Rightarrow scission point)
- Only a very few relevant deformation parameters

$$\varrho_s^2(z) = \mathcal{N} (1 - u^2) \left[1 - B e^{-\beta(u-\alpha)^2} \right] \quad u = \frac{z - z_{sh}}{z_0} \quad z_0 = c R_0$$

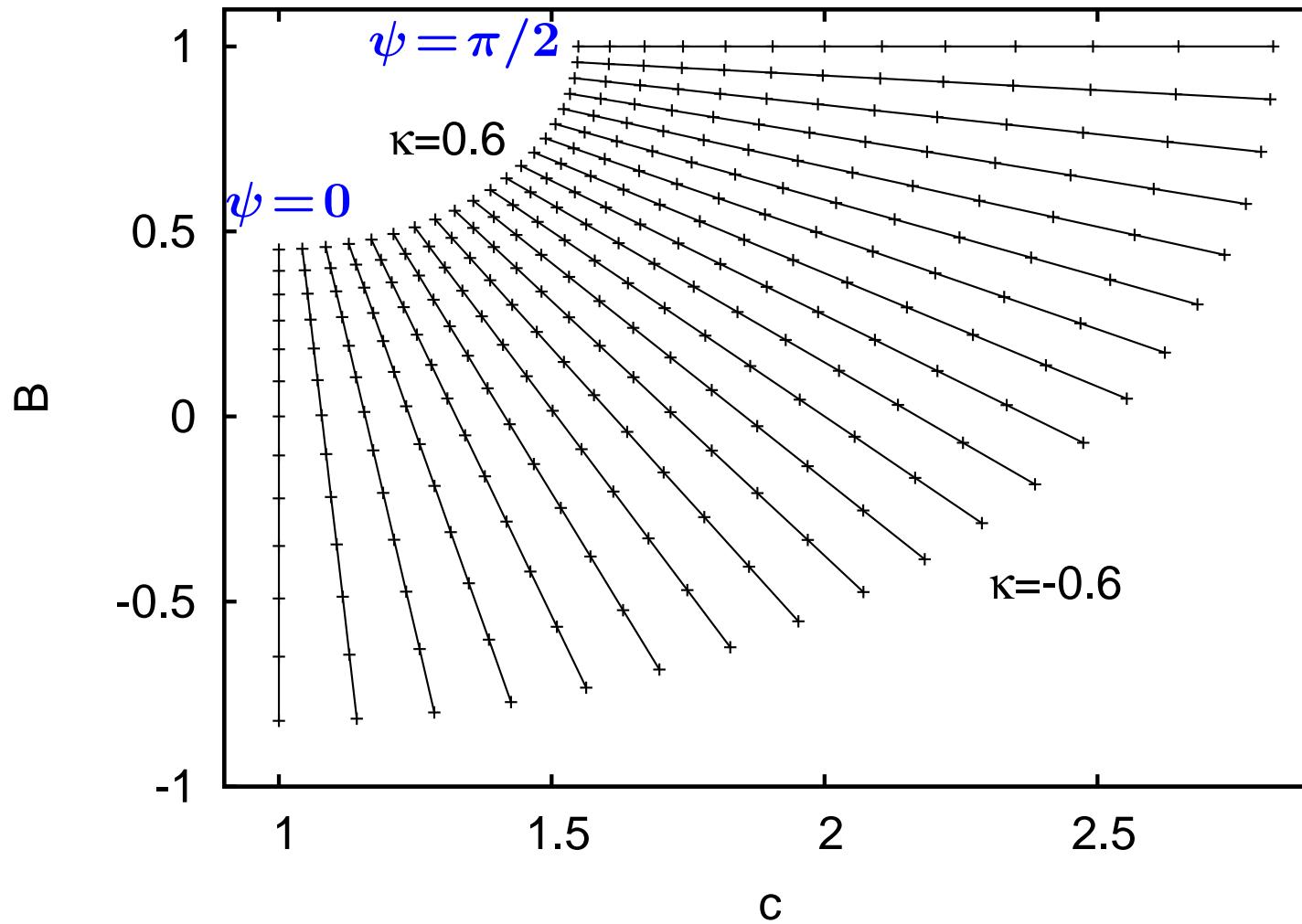
\mathcal{N} normalization constant that depends on c , B and α



Difficulty: localize precisely the scission configuration

Solution: introduce new parameters ψ and κ

$$c = 1 + e^{-\kappa} \sin \psi \quad \text{and} \quad B = 1 - e^{-\kappa} \cos \psi$$



breaking axial symmetry

suppose ellipsoidal shape \perp to z axis
and introduce the non-axiality parameter

$$\eta = \frac{a_y - a_x}{a_y + a_x}$$

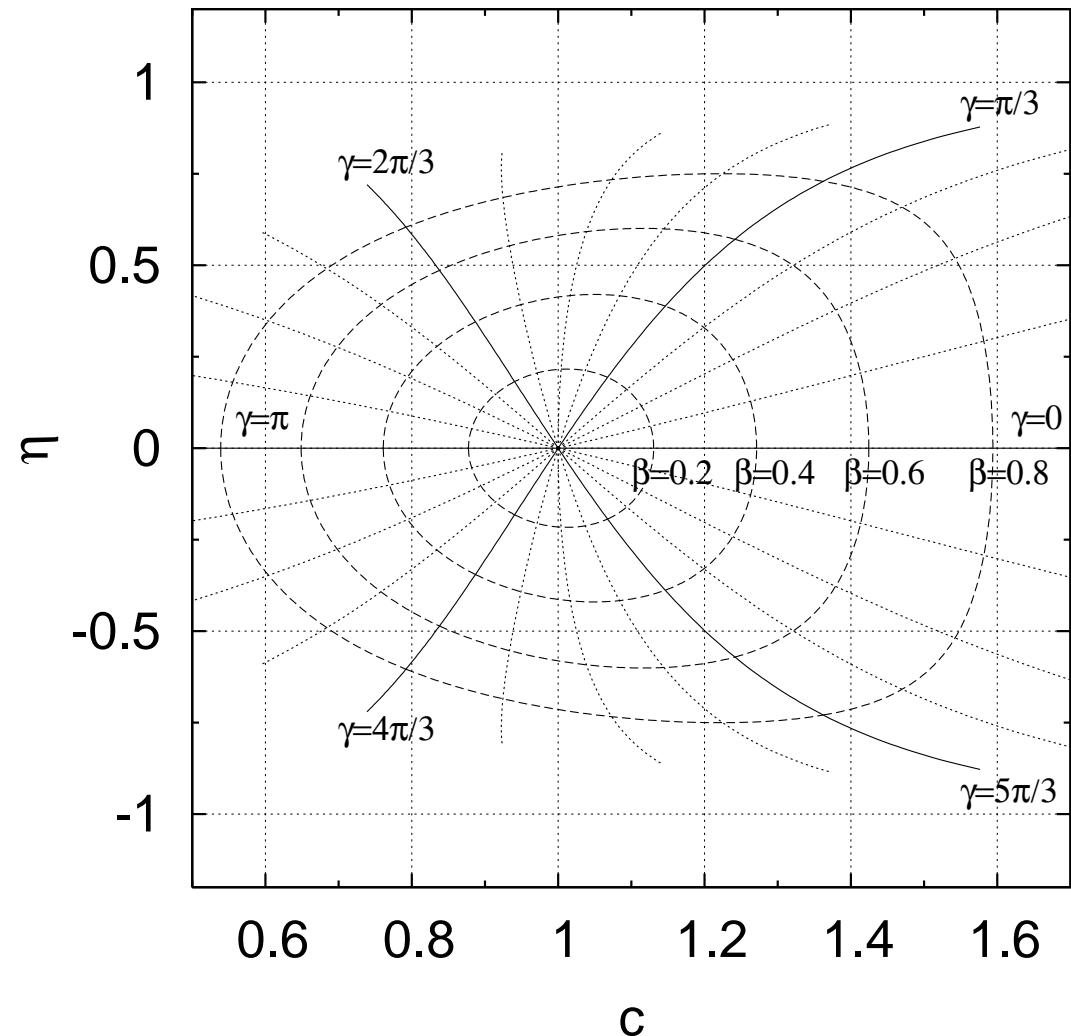
assume that η is independent of z

$$a_x(z) = \varrho_s(z) \left(\frac{1 - \eta}{1 + \eta} \right)^{1/2}$$

$$a_y(z) = \varrho_s(z) \left(\frac{1 + \eta}{1 - \eta} \right)^{1/2}$$

volume conservation then leads to

$$\tilde{\varrho}_s^2(z, \varphi) = \varrho_s^2(z) \frac{1 - \eta^2}{1 + \eta^2 + 2\eta \cos(2\varphi)}$$



Transformation from (c, η) to (β, γ) in the pure spheroidal case

Lublin-Strasbourg Drop Model

According to the Strutinski theorem

$$E_{\text{tot}} = \int \mathcal{H}(\rho) d^3r = E_{\text{mac}} + \delta E_{\text{mic}}$$

with the Lublin-Strasbourg drop

$$\begin{aligned} E_{\text{mac}}(Z, N, def) &= a_{\text{vol}} (1 - \kappa_{\text{vol}} I^2) A + a_{\text{surf}} (1 - \kappa_{\text{surf}} I^2) A^{2/3} B_{\text{surf}}(def) \\ &\quad + a_{\text{cur}} (1 - \kappa_{\text{cur}} I^2) A^{1/3} B_{\text{cur}}(def) + \frac{3e^2}{5r_o} \frac{Z^2}{A^{1/3}} B_{\text{Coul}}(def) \\ &\quad - C_4 \frac{Z^2}{A} - E_{\text{cong}} \end{aligned}$$

with shell and pairing corrections

$$\delta E_{\text{shell}} = \sum_{\text{occ}} \varepsilon_\nu - \tilde{E}$$

$$\delta E_{\text{mic}} = \delta E_{\text{shell}} + \delta E_{\text{pair}}$$

$$\delta E_{\text{pair}} = E_{\text{BCS}} - \sum_{\text{occ}} \varepsilon_\nu - \langle E_{\text{pair}} \rangle$$

Energy landscapes in light + medium-mass nuclei

studied nuclei: ^{46}Ti , ^{92}Mo , ^{138}Ba

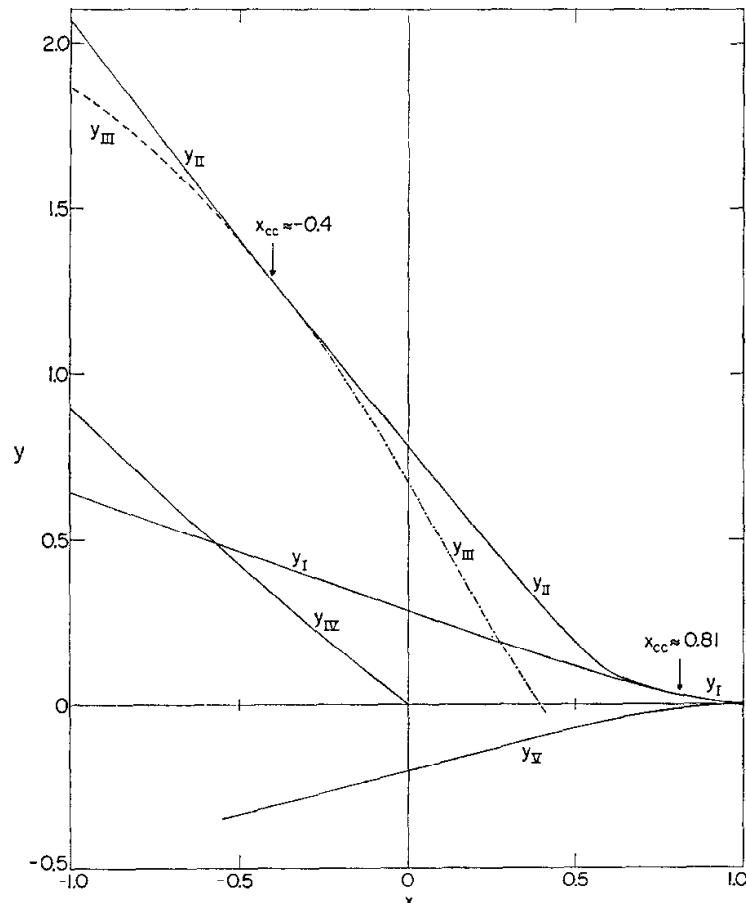
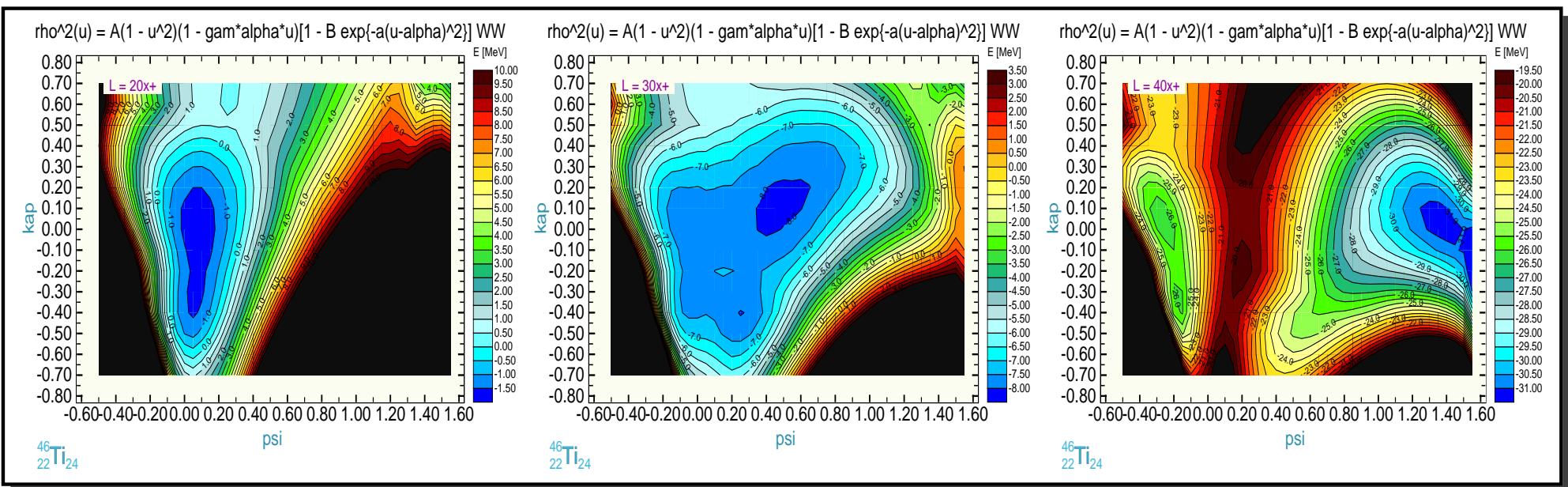
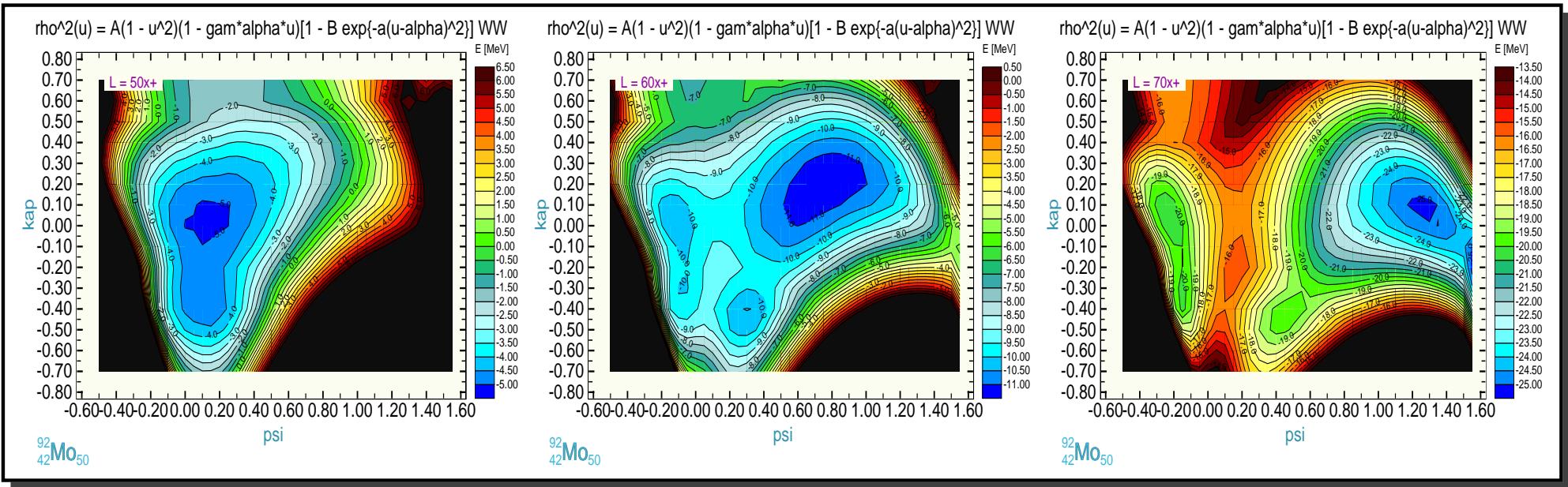


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Instability against
reflexion asymmetry
when $x \leq 0.4$ for $L=0$,
decreasing for $L>0$

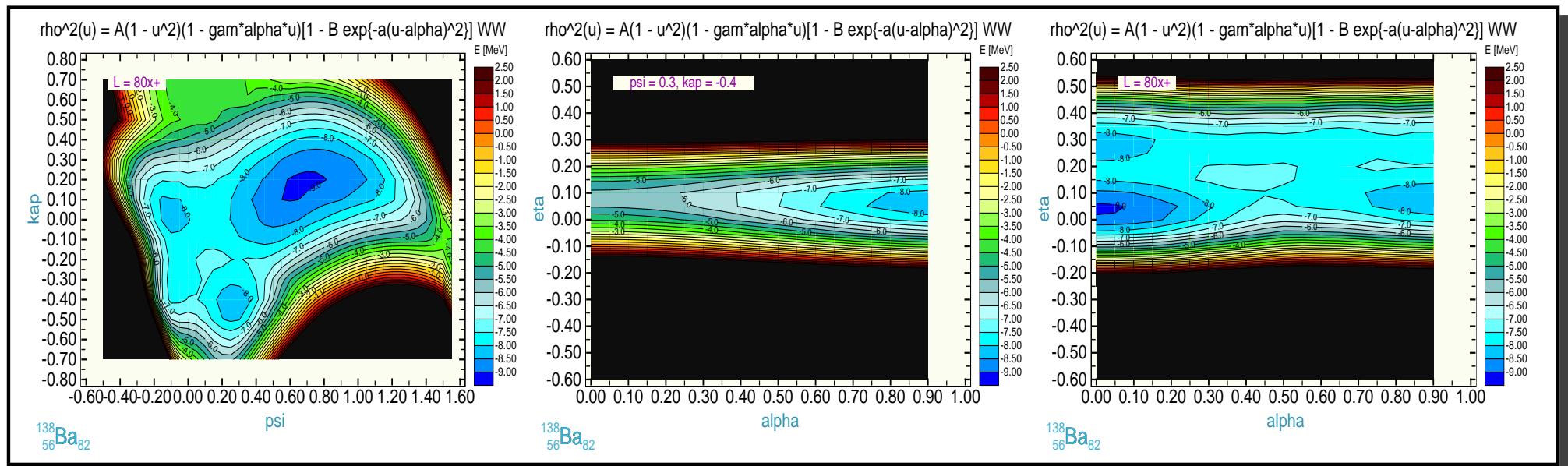
A rough estimate:

$$x \approx \frac{1}{50} \frac{Z^2}{A}, \quad y \approx 2 \frac{L^2}{A^{7/3}}$$



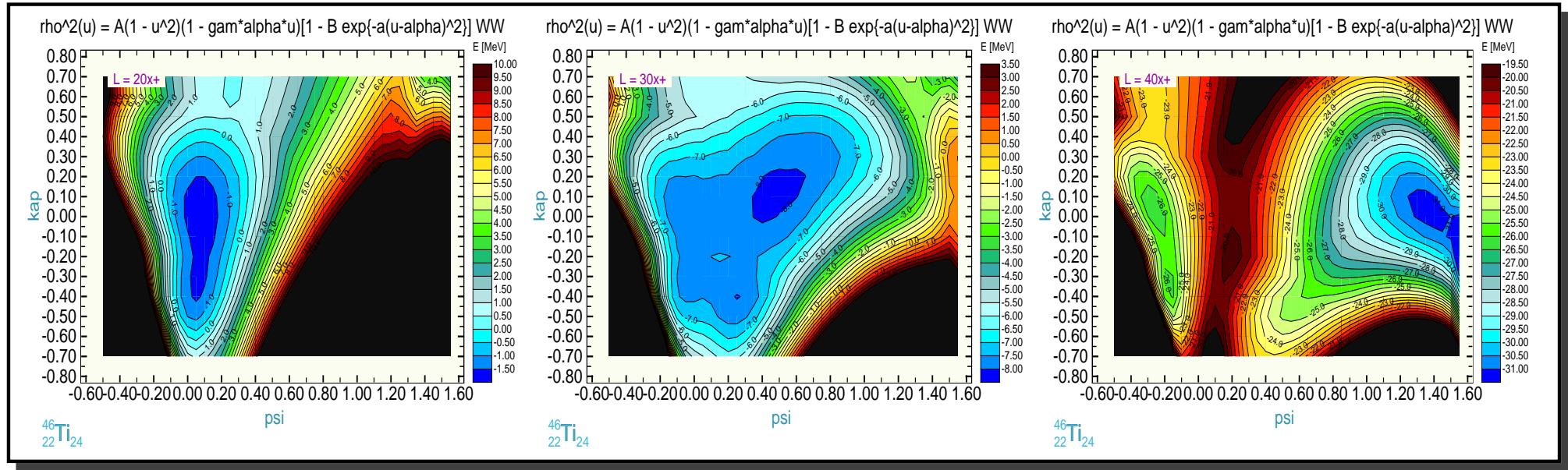
The Poincaré instability

also motivated by the experimental studies of A. Maj & coworkers
 (see also talk by Kasia Mazurek this conference)



in ^{138}Ba

in a light nucleus as a function of L



^{46}Ti

Remark:

rotation stabilizes the nucleus versus the Poincaré instability
at fixed elongation

can we understand that?

consider moment of inertia of 2 touching spheres ($\alpha = 0$) $\mathcal{J} \downarrow$
with one of them swallowing the other one ($\alpha \rightarrow 1$)

$L \uparrow$ how to get more stability? $\Rightarrow E_{\text{rot}} \downarrow$ $\mathcal{J} \uparrow \Rightarrow \alpha \downarrow$

Conclusions

- Modified Funny-Hills (MFH) parametrisation describes nuclear shapes very well (already well known)
- Our LDM study with the LSD and the MFH parametrization allows for a systematic study of the Poincaré instability in light nuclei
- This instability against left-right asymmetry for certain stationary points already present in medium-mass nuclei with mass $A \approx 130$
- Higher angular momenta tend to stabilize the nucleus already pointed out by CPS 35 years ago
- More nuclear systems need to be studied, in particular concerning the effect of finite temperatures
- Comparison with Fedir Ivanyuk's result rather difficult
“optimal” fission valley \Leftrightarrow 4D energy landscape