COMPETITION OF ALPHA DECAY AND HEAVY PARTICLE DECAY IN SUPERHEAVY NUCLEI

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OUTLINE

- Macroscopic-microscopic method
- Unified approach of cold fission, α-decay and heavy particle radioactivities (HPR) within ASAF model
- Experimental confirmations
- New mass table Audi & Meng. KTUY05 and FRDM95
- α -decay and HPR of heaviest superheavies
- Results within ASAF, UNIV and semFIS
- Summary



Macroscopic-microscopic method

Accounting for quantum single-particle structure and classical collective properties.

Liquid Drop Model: E_{LD}

Single-particle shell model (SPSM): energy levels vs. deformation. *Two-center shell model for fission and fusion.*

Shell correction method: $\delta E = \delta U + \delta P$

Total deformation energy: $E_{def} = E_{LD} + \delta E$

The potential of SPSM Hamiltonian should admit the drop eq. $\rho = \rho(z)$ as an equipotential surface. Semi-spheroidal shape, allows to obtain analytical results for atomic clusters on a surface.



Intersected spheres



Two intersected spheres. Volume conservation and $R_2 = \text{const.}$ One deformation parameter: separation distance R. Surface equation $\rho = \rho(z)$. Initial $R_i = R_0 - R_2$. Touching point $R_t = R_1 + R_2$.

 $(R-R_i)/(R_t-R_i)=0.25$ Example: $^{232}U \rightarrow ^{24}Ne + ^{208}Pb$ 1.0 0.50 0.75 0.8 · 1 25 Two center shell model (Frankfurt) potential 0.6 ⁰/^z/^z/ 0.4 $(R-R_i)/R_{ti}=0$ 0.25 0.50 0.75 0.2 0.0 -0.5 0.0 0.5 1.0 1.5 2.0 -1.0 Sequence of shapes z/R_0



Liquid drop model

Nucleus considered a uniformly charged drop. Two variants: LDM and Yukawa-plus-exponential (Y+EM). LDM (surface + Coulomb) deformation energy

 $E_{LDM} = E - E^{0} = (E_{s} - E_{s}^{0}) + (E_{C} - E_{C}^{0})$ $= E_{s}^{0}(B_{s} - 1) + E_{C}^{0}(B_{C} - 1)$ For spherical shapes $E_{s}^{0} = a_{s}(1 - \kappa I^{2})A^{2/3}$; I = (N - Z)/A; $E_{C}^{0} = a_{c}Z^{2}A^{-1/3}$. Nuclear fissility $X = E_{c}^{0}/(2E_{s}^{0})$.

Parameters obtained by fit to experimental data on nuclear masses, quadrupole moments and fission barriers: $a_s = 17.9439$ MeV, $\kappa = 1.7826$, $a_c = 3e^2/(5r_0)$, $e^2 = 1.44$ MeV·fm, $r_0 = 1.2249$ fm. W.D. Myers and W.J. Swiatecki, Nucl. Phys. A 81 (1966) 1



Shell corrections

The total energy of the uniform level distribution

$$\tilde{u} = \tilde{U}/\hbar\omega_0^0 = 2\int_{-\infty}^{\lambda} \tilde{g}(\epsilon)\epsilon d\epsilon$$

In units of $\hbar \omega_0^0$ the shell corrections are calculated for each deformation ε

$$\delta u(n,\varepsilon) = \sum_{i=1}^{n} 2\epsilon_i(\varepsilon) - \tilde{u}(n,\varepsilon)$$

 $n = N_p/2$ particles. Then $\delta u = \delta u_p + \delta u_n$.



Pairing corrections

The gap Δ and Fermi energy λ are solutions of the BCS eqs:

$$0 = \sum_{k_i}^{k_f} \frac{\epsilon_k - \lambda}{\sqrt{(\epsilon_k - \lambda)^2 + \Delta^2}} \quad ; \quad \frac{2}{G} = \sum_{k_i}^{k_f} \frac{1}{\sqrt{(\epsilon_k - \lambda)^2 + \Delta^2}}$$

$$k_i = Z/2 - n + 1, \quad k_f = Z/2 + n', \quad \frac{2}{G} \simeq 2\tilde{g}(\tilde{\lambda}) \ln\left(\frac{2\Omega}{\tilde{\Delta}}\right).$$

The pairing correction $\delta p = p - \tilde{p}$, represents the difference between the pairing correlation energies for the discrete level distribution $p = \sum_{k=k_i}^{k_f} 2v_k^2 \epsilon_k - 2\sum_{k=k_i}^{Z/2} \epsilon_k - \frac{\Delta^2}{G}$ and for the continuous level distribution $\tilde{p} = -(\tilde{g}\tilde{\Delta}^2)/2 = -(\tilde{g}_s\tilde{\Delta}^2)/4$. Compared to shell correction, the pairing correction is out of phase and smaller. One has again $\delta p = \delta p_p + \delta p_n$, and $\delta e = \delta u + \delta p$.



Example: Na $_{148}$ atomic cluster



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 $E_v = -333 \text{ eV}$ was not included in E_{LD} and E. Liquid drop and total deformation energy (top). Shell plus pairing corrections for hemispheroidal harmonic oscillator energy levels (bottom). Smoothing effect of pairing. Ground state shape prolate $\delta = 0.47$ Semiaxes ratio $\frac{a}{c} = \frac{2-\delta}{2+\delta}$

²²²**Ra** E_{Y+EM} , $\delta E_{shell+pair}$, E_{def} **PES**







separation distance $\xi = (R - R_i)/(R_t - R_i)$ mass asymmetry $\eta = (A_1 - A_2)/(A_1 + A_2)$

Poenaru, Gherghescu, W.Greiner, Phys. Rev. C 73 (2006) 014608



Basic relationships

Parent \rightarrow emitted ion + daughter nucleus, ${}^{A}Z \rightarrow {}^{A_{e}}Z_{e} + {}^{A_{d}}Z_{d}$ Measurable quantities

Kinetic energy of the emitted cluster $E_k = QA_1/A$ or the released energy $Q = M - (M_e + M_d) > 0$.

Decay constant $\lambda = \ln 2/T$ or Half-life ($T < 10^{32}$ s) or branching ratio $b_{\alpha} = T_{\alpha}/T$ ($b_{\alpha} > 10^{-17}$)

Model dependent quantities ($\lambda = \nu SP_s$)

- $\checkmark \nu$ frequency of assaults or $E_v = h\nu/2$
 - S preformation probability
- P_s penetrability of external barrier



Fission theory

Shape parameters: fragment separation, R, and mass asymetry $\eta = (A_d - A_e)/A$. Our method to estimate preformation as penetrability of internal

Our method to estimate preformation as penetrability of internal barrier: $S = \exp(-K_{ov})$. DNP, WG, *Physica Scripta* 44 (1991) 427. Similarly $P = \exp(-K_s)$ for external barrier. Action integral calculated within Wentzel-Kramers-Brillouin (WKB) quasiclasical approximation

$$K_{ov} = \frac{2}{\hbar} \int_{R_i}^{R_t} \sqrt{2B(R)E(R)} dR$$

E - Potential barrier

 $B = \mu$ – Nuclear inertia = reduced mass for $R \ge R_t$



Unified approach: CF; HPR, and α **-d**







Three valleys: cold-fission (almost symmetrical); ¹⁶O radioactivity, and α -decay

²³⁴U half-lives spectrum (short T up)



Experimental confirmations

Rare events in a strong background of α particles Detectors:

- Semiconductor telescope + electronics
- Magnetic spectrometers (SOLENO, Enge split-pole)
- Solid state nuclear track det. (SSNTD). Cheap and handy. Need to be chemically etched then follows microscope scanning

Experiments performed in Universities and Research Institutes from: Oxford; Moscow; Orsay; Berkeley; Dubna; Argonne; Livermore; Geneva; Milano; Vienna, and Beijing. Table: R. Bonetti and A. Guglielmetti, Rom. Rep. Phys. **59** (2007) 301.



Systematics $T_{1/2}$: ¹⁴C, ^{18,20}O, ²³F rad.



new confirm — A. Guglielmetti et al., J Phys: Conf Ser 111 (2008) 012050 One of the new candidates from our paper: Poenaru, Nagame, Gherghescu, W. Greiner *Phys. Rev.* C 65 (2002) 054308.



Systematics $T_{1/2}$: ^{22,24,25,26}**Ne rad.**



Only lower limits for ¹⁸O and ²⁶Ne



Systematics $T_{1/2}$: ^{28,30}**Mg**, ^{32,34}**Si rad**.



CLUSTER Decays Dorin N. POENARU, IFIN-HH

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New table of experimental masses



3290 nuclei,
2377 measured
and 913 det. from
Systematics.
G. Audi, W.
Meng, Private
communication
2011.

C6 CLUSTER DECAYS

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18th Nuclear Physics Workshop Maria and Pierre Curie, 28 Sept - 02 Oct 2011, Kazimierz Dolny, Poland – p.17/34

KTUY05 Calculated Masses



9441 nuclei with Z=2-130 and N=2-200. H. Koura, T. Tachibana, M. Uno and M. Yamada, *Prog. Theor. Phys.* **113** (2005) 305.



FRDM95 Calculated Masses



8979 nuclei with Z=8-136 and N=8-236. P. Möller, J.R. Nix, W.D. Myers, W.J. Swiatecki, *At. Data Nucl.Data Tables* **59** (1995) 185.



Examples of time spectra (I)





Examples of time spectra (II)



 $^{288}114 \rightarrow {}^{80}Ge + {}^{208}$ Pb. D.N. Poenaru, R.A. Gherghescu, W. Greiner, *Phys. Rev. Lett.* **107** (2011) 062503.



Superheavies as cluster emitters



New concept: for $Z > 110 Z_e > 28$ to get a daughter around ²⁰⁸Pb. D.N. Poenaru, R.A. Gherghescu, W. Greiner, *Phys. Rev. Lett.* **107** (2011) 062503.



HIR of SH nuclei (I)



Half-lives

Branching ratios

Branching ratio with respect to α decay: $b_{\alpha} = T_{\alpha}/T_c$. Usually $b_{\alpha} << 1$. Trend: shorter T_c and larger b_{α} . For larger Z > 120 there are SHs with $T_c < 1$ ns and $b_{\alpha} > 1$.



HIR of SH nuclei (II)



KTUY05

FRDM95



HIR of SH nuclei (III)



KTUY05

FRDM95



SH nuclei as cluster emitters



FRDM95 $Z_e \leq Z - 80$ (freq. daughter around ²⁰⁸Pb) Most probable emitted clusters with different colors.



Universal curves (I)

Approximations: $\log S = [(A_e - 1)/3] \log S_{\alpha}$, $\nu(A_e, Z_e, A_d, Z_d) = \text{constant. From fit to } \alpha \text{ decay:}$ $S_{\alpha} = 0.0160694 \text{ and } \nu = 10^{22.01} \text{ s}^{-1}.$

 $\log T = -\log P - 22.169 + 0.598(A_e - 1)$

$$-\log P = c_{AZ} \left[\arccos \sqrt{r} - \sqrt{r(1-r)} \right]$$

$$c_{AZ} = 0.22873(\mu_A Z_d Z_e R_b)^{1/2}, \ r = R_t/R_b, \ R_t = 1.2249(A_d^{1/3} + A_e^{1/3}), \ R_b = 1.43998Z_d Z_e/Q, \text{ and } \mu_A = A_d A_e/A.$$
DN Poenaru, W Greiner, *Physica Scripta* **44** (1991) 427.



Universal curves (II)



Geiger-Nuttal plot $T_{\alpha} = f(\text{range of } \alpha \text{ in air})$ $\log T = f(1/Q^{-1/2})$



Single Universal curve: α **and HIR**



D.N. Poenaru, R.A. Gherghescu, W. Greiner, Phys. Rev. C, 83 (2011) 014601.



α -decay Z = 92 - 118, fission models



Vertical bars: $N_d = 126, 152, 162$ Poenaru, D.N., Plonski, I.H., and Greiner, W., *Phys. Rev.* C 74 (2006) 014312



$\alpha\text{-decay:}$ ASAF, semFIS & UNIV



STANDARD DEVIATION

Group	σ -ASAF	σ -UNIV	σ -semFIS
47 e-e	0.402	0.267	0.164
45 e-o	0.615	0.554	0.507
25 о-е	0.761	0.543	0.485
25 0-0	0.795	0.456	0.451

Poenaru, D.N., Plonski, I.H., Gherghescu, R.A., Greiner, W., *J. Phys. G* **32** (2006) 1223



UNIV and UDL curve for e-e nuclei



Vertical bars: N = 50, 82, 126, 152, 162, 172UDL: C. Qi, F.R. Xu, R.J. Liotta, R. Wyss Phys. Rev. Lett. **103** (2009) 072501.



Compare UNIV, UDL, semFIS for α **decay**

Standard deviations of log T values of semFIS, UNIV and UDL formulae for 173 even-even, 134 even-odd, 123 odd-even, and 104 odd-odd α emitters with atomic numbers Z = 52 - 118.

n	σ_{UNIV}	h_{UNIV}	σ_{UDL}	c_{UDL}	σ_{SEM}
173	0.354	0.030	0.329	-20.796	0.222
134	0.640	0.528	0.606	-20.327	0.501
123	0.565	0.379	0.538	-20.481	0.434
104	0.826	0.961	0.804	-19.904	0.567

$$\sigma = \left\{ \sum_{i=1}^{n} [\log(T_i/T_{exp})]^2 / (n-1) \right\}^{1/2}$$



Summary

- ASAF model predictions have been confirmed for parent nuclei with Z = 87 96
- The magicity of the daughter ²⁰⁸Pb was not fully exploited: new experimental searches can be performed
- The ASAF, semFIS and universal curves UNIV and UDL provide good estimation of half-lives
- For some superheavies HIR half-lives could be shorter than that of α decay

