Fusion hindrance and SHE

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On the prediction for the SHE production

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Questions for theoreticians

Can we guide the experiments ?

What is the shell correction energy ?

Difficulties

Models cannot be extrapolated from lighter systems

+ Fusion hindrance

Extremely low cross sections

+ Few data with few information



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Direct evidence for long fission times



M. Morjean et al, Eur. Phys. J. D45 (2007) 27 & PRL101, 072701 (2008) Mesurements at GANIL by crystal blocking techniques

Z = 124 A = 312

At least 12 % of the capture events with a life time longer than 10^{-18} s

Z = 120 A = 296

At least 10 % of the capture events with a life time longer than 10^{-18} s

Z = 114 A = 282

Very low statistic or no events with a life time longer than 10⁻¹⁸ s

 $\Gamma_{f} \approx 10^{-21} s$

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KEWPIE 2

Γ_{n}



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 $\frac{dP_i}{dt} = \Gamma_{i-1}^n P_{i-1}(t) - (\Gamma_i^f + \Gamma_i^n) P_i(t)$

Main ingredients

- Formalism:
 - Weisskopf or Hauser-Feshbach
 - Bohr-Wheeler
- Level density:
 - Bohr-Mottelson with angular dependence
- Level density parameter:
 - Töke-Swiatecki
 - Suppression of shell energy corrections according to lgnatyuk
- Collective enhancement included

A. Marchix, PhD thesis, Univ. Caen (2007)

Simplified model

Fission vs neutron evaporation

Bn=6 MeV & Bf

chain

mean:

 $\bullet B_f \approx B_n$

constant along the

Long fission times

without gamma emission - with gamma emission **Average fission time (s)** 1E-15 1E-16 1E-17 1E-18 1E-19 1E-20 1E-16 1E-19 8 **Pre-scission neutrons** n neutron multiplicity 7 9 2 4 E*=80 MeV 3

Long fission time events occur after evaporation of several neutrons

♦ We cannot extract Bf of each isotope

D.B. et al, IJMP E17 (2008) 1681-1693

0

0

0,5

2

1,5

Bf/Bn





- Experimental results for Z=120 and 124 cannot be reproduced with Möller's table
- Very large △E_{shell} for the first isotopes of the evaporation chain
 - Potential structure effects ?

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Structure effect 2



Residue cross sections



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Specificity

- It is not a Monte-Carlo code to calculate very low probabilities
- It is based on a discretisation in bins of the energy spectra:



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Residue cross sections

Important parameters of KEWPIE2

Shell correction energy -> correction factor

$$\Delta E_{shell} = f \cdot \Delta E_{Moller}$$

Damping Energy E_d

$$a_{ground} = a.(1 + \frac{(1 - e^{-E^*/E_d}) \Delta E_{shell}}{E^*})$$

• Originally, Ed=18.5 MeV

$$B_f \approx e^{-E^*/E_d} \Delta E_{shell}$$

Reduced friction

β=2.10²¹s⁻¹

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■ Fitting the residue cross sections gives very strong constraint on ∆E_{shell}...

Precision of 1 MeV

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... if we know the fusion cross section





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Experimental fusion hindrance



K.-H. Schmidt & W. Morawek Rep. Prog. Phys. 54 (1991) 949



Position of the inner barrier





Neck dynamics



Fusion hindrance for symmetric reactions



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Fusion hindrance for symmetric reactions



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Influence of the shift



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Borderline between hindered and non hindered reactions



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Adiabatic approximation

- Fast degree of freedom (neck) vs slow degree of freedom (r)
- Tensor coupling
- Shift of the initial condition in r
- Larger hindrance due to dynamical coupling!



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